

of highly trained and experienced geologists than to that of gentlemen, who, whatever their attainments may be, certainly do not in this Report evince much acquaintance with geology.

J. G.

Reports of the Mining Surveyors and Registrars for Quarter ending March 31, 1872. Victoria.

THERE is nothing in these Reports calling for special notice. The total quantity of gold got respectively from alluvia (or, as the Reports have it, *alluviums*) and quartz reefs during the quarter were as follows:—Alluvial, 171,851 oz. 10 dwt.; quartz, 164,670 oz. 8 dwt.; total, 336,521 oz. 18 dwt. The quantity of gold, the produce of the colony, exported during the same period was 398,131 oz. 10 dwt.

LETTERS TO THE EDITOR

[The Editor does not hold himself responsible for opinions expressed by his correspondents. No notice is taken of anonymous communications.]

Aurora Spectrum

IN connection with my letter in last number of NATURE, I have in a diagram approximately placed the aurora lines side by side with the spectrum of hydrogen and of some of the principal air lines (as given in Dr. Watt's index) and with the following results:—

Line No. 1. Close upon, if not identical with, an air-line marked by Huggins N O, and Plücher O.

No. 2.—Not apparently coincident with any prominent air-line. The coincidence with a line of oxygen noted by so careful an observer as Mr. Proctor is puzzling; and if the instrumental power used was sufficient to ensure absolute identity, seems to indicate a second or unusual spectrum of that gas.

No. 3 is not near any principal air-line.

No. 4 is nearly coincident with a faint line of oxygen (confirmatory of Mr. Proctor's observations).

No. 5 corresponds to a rather strong N line.

No. 6 does not coincide with any principal air-line, very faint lines of O and N being the nearest.

No. 7. Upon close examination the positions of this line as respectively fixed by Mr. Proctor and Lord Lindsay are not inconsistent, and the line closely corresponds with a strong line of oxygen situated on the less refrangible side of solar G.

Nos. 1, 4, 6, and 7 fairly correspond in intensity with their representative air-lines. None of the lines are identified with H α , H β , or H γ , and it would appear that the aurora, if a spectra of atmospheric gases, mainly selects oxygen and ignores H α and the stronger N lines. The modification of compound spectra by conditions of temperature and pressure, is however only a partially explored subject, and we have moreover no certain data of conditions in the case of the aurora, which will assist us in bringing it to bear.

I accidentally omitted from the names of some observers of the zodiacal light that of Prof. Piazz Smyth, whose observations in the south may be said to have conclusively demolished the supposed identity of the light, and the aurora (at least so far as bright lines are concerned) made it extremely improbable that anything beyond a continuous spectrum will ever be seen in the pure zodiacal light, though a further search should be by no means neglected.

J. RAND CAPRON

Guildford, Jan. 10, 1873

Polarisation of the Zodiacal Light and of the Aurora

IN the interesting article by Mr. Rand Capron in the last number of NATURE—after collating the various results of the spectroscopic examinations of the aurora and zodiacal light which have appeared at different times in your pages, together with those which have been collected by Dr. Schellen—he terminates his analysis of the general results by remarking that he is “not aware whether the zodiacal light and the aurora have been examined with the polariscope,” and suggests that the “light, though faint, might be tested with a Nicol's prism and Savart's bands.”

I would refer him to a paper in the March number of the “Monthly Notices of the Astronomical Society” for 1871, in

which an observation by Mr. Burton (late assistant to the Earl of Rosse) on the polarisation of the zodiacal light is described.

Mr. Burton was one of the eclipse party stationed at Agosta, in Sicily. He made use of a Savart's polariscope, set so as to give a black centre where the bands were parallel to the plane of polarisation. On looking to the brightest parts of the zodiacal light Mr. Burton believed that he could detect faint traces of polarisation, sufficiently strong to enable him just to recognise that the bands were black centred when their direction coincided with the axis of the cone of light, that is, when the direction passed through the position of the sun.

To make sure that he was not examining the remains of air polarisation given by the slight remaining twilight, he examined the light of other parts of the heavens, but was unable anywhere else to detect any trace of bands. In contradistinction, however, to this must be set an observation of my own, yielding a negative result, though made on the same evening and with a similar instrument, as well as with the same Savart used by Mr. Burton. I was, however, unable to detect any trace of bands either upon the cones of zodiacal brightness or upon the adjacent parts of the sky; but it is very possible that Mr. Burton's eye may be more sensitive to faint lights than my own.

In February last I also met with a negative result in examining a faint trace of the zodiacal light visible in England. I then used a double-image prism as well as a Savart, thinking that its two oppositely polarised fields in juxtaposition might afford a more delicate test for so faint an object.

Capt. Tupman while cruising in the Mediterranean has also, I believe, repeatedly obtained negative results when making use of a Savart on the zodiacal light.

And I understand that Mr. Lockyer, together with the other observers of the Indian Eclipse of December 1871, totally failed to detect any traces of polarisation in the brilliant displays of the zodiacal light which they observed while crossing the Indian Ocean.

I am therefore disposed to conclude that any traces of polarisation must be very slight, if indeed any polarisation at all is to be attributed to the zodiacal light itself and not to the veil of atmospheric impurities lying between us and objects near to the horizon. Certainly we may conclude that there is no such polarisation as is found in the light of the solar corona or—as we might expect—if the zodiacal light were caused by a great cloud of cosmical dust made up of particles smaller in diameter than the wave-length.

Indeed there cannot be as great a percentage of polarisation, or, to speak more exactly, as great a difference between the component radial to the sun's place and the component at right angles, as in the case of a sunbeam dispersed by the dust in our own atmosphere. For if any one will examine the track of a sunbeam passing through a room with a Savart, he will not fail to be struck with the distinctness of the bands. We seem therefore justified in concluding, that if the zodiacal light is composed of cosmical dust, such dust particles must be considerably coarser than those which float in our own atmosphere.

As to the polarisation of the light of the aurora, I examined, both with a double-image prism and Savart, a faint auroral display on November 10, 1871; as also the light of the great aurora of Sunday, Feb. 4, 1872, but in neither instance was able to detect any traces of polarisation.

A. COWPER RANYARD

The Diathermancy of Flame

THERE are some statements in Capt. Ericsson's reply to my letter (NATURE, vol. vii. p. 149) which demand discussion. In the first place he calculates the supply of gas in *his* pipe and applies it to *my* burners. As his pipe did not supply my flames, but his own, which were at least fifteen times larger than mine, the applicability of his figures is rather obscure.

Capt. Ericsson says, “The apparatus contrived by Mr. Williams for determining the diathermancy of flame, as described by himself, is exceedingly faulty, the temperature it records being that produced by heat derived from several sources. The radiant heat transmitted to the bulb of the thermometer by the flames, acting conjointly with the unknown degree of heat imparted by the surrounding medium, it will be evident that Mr. Williams' device is worthless as an indicator of radiant intensity.” Does Capt. Ericsson really mean that the maximum temperature indicated by a thermometer exposed to several varying sources of heat is not determined by the maximum radiators or convections

of the body capable of communicating the highest temperature, but by this, plus the minor radiators or convections of the cooler bodies? The words I have put in italics distinctly imply such an assumption.

He seems to forget that I did, in the first place, observe and record the temperature of the surrounding medium. It was the 19° C. which served as my starting-point. As no additional radiations were introduced beyond those of the flames to be experimented upon, and the blackened bulb of my thermometer was surrounded by polished reflecting metal surfaces on all sides, except that exposed to the flames, all the subsequent increments of heat were unquestionably due to the radiators from those flames, whether they came directly from the flames themselves or were received and reflected from the back and sides of the polished chamber. Fully admitting the desirability of a continuous record of the heat thus communicated to the surroundings of the thermometer during the experiments, I nevertheless firmly maintain that, rude as it was, my apparatus (I refer, not to the thermometer, but to its adjuncts) was far superior to Capt. Ericsson's. Mine was liable to a small source of error from a possible accidental irregularity of radiation by the thermometer bulb, but his was specially devised to ensure a large amount of such irregularity, continually increasing with the progress of the experiments. It is not a little surprising that so careful and luxurious an experimentalist as Capt. Ericsson should have overlooked the fact, that the very precautions which he so elaborately introduced to secure equal radiation from his bulb are precisely adapted to produce the contrary result.

The arrangements by which his thermometer is "enclosed in an exterior vessel charged with water kept at a constant temperature" of 60° by communication with a capacious cistern, directly violate the conditions demanded by the Newtonian law of radiation, of which Capt. Ericsson is so able a champion; for as the experiment proceeds with an increasing number of flames, and consequent rising of the thermometer, this constant temperature of the water jacket goes on steadily augmenting the difference between the temperature of the bulb and that of its surroundings, and consequently secures just what it is intended to prevent, viz. a variable radiation. What is required to secure a constant degree of radiation from the bulb is not the constant temperature of the surroundings, but a temperature steadily increasing at the same rate as that of the bulb, in order that the *differential* and not the *absolute* temperature of the surrounding medium, &c., should remain constant. This was rudely obtained in my simple apparatus, as both the thermometer and its surroundings were simultaneously influenced by the same radiations.

Capt. Ericsson takes great pains to controvert my "assumption that the intensity of a gas flame is proportional to the gas consumed." This is unnecessary, inasmuch as I never made any such assumption, but have, on the contrary, endeavoured to prove that such cannot possibly be the case, by showing what becomes of the radiations from the interior of a large solid flame. If he will read chaps. 7 and 8 of "The Fuel of the Sun," he will see how and why this has been done, and learn the true bearings of the experiments under discussion upon this subject.

W. MATTIEU WILLIAMS

P.S.—The present is a suitable opportunity for asking a question which doubtless the philological readers of *NATURE* can easily answer. Many writers use the words "diathermancy," "diathermanous," "athermanous," &c., rather than "diathermacy," "diathermous," &c. Why is this? We do not say "thermanal" or "thermanometer," &c. Why, then, should we depart from the analogy of ancient usage in constructing the more modern compounds of the same root?

Pollen-eaters

MR. HART's note in *NATURE*, vol. vii. p. 161, is interesting to those who have paid attention to the subject of fertilisation by insect agency, and would be still more so if he could furnish the names of the species of both plants and *Syrphidae* that have come under his observation.

May I take this opportunity of calling the attention of the readers of *NATURE* to a suggestion which I made some months since in the *Journal of Botany*, and which has at present met with no response? I believe no greater service could be rendered to this department of physiological botany than a series of observations on the species of insects which frequent and assist in the fertilisation of our wild flowers. I know of no such list even

with respect to our commonest flowers. Here is a wide field for observation during the next season.

London, Jan. 7

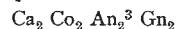
ALFRED W. BENNETT

P.S.—At the time of writing the above, I had not seen Dr. Buchanan White's article in the January number of the "Journal of Botany," on "The Influence of Insect-agency in the Distribution of Plants," an admirable introduction to the series of papers I had in my mind.

Welwitschia

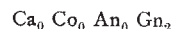
If you will kindly permit me, I wish to make an addition to your notice of my paper on "Welwitschia," read at the Linnean Society on the 19th ult. That paper was completed and put in Dr. Hooker's hands about three months ago; and the reading of it was delayed until I had seen Strasburger's recently published memoir on *Coniferae* and *Gnetaceae*. After perusing that valuable work, I added a small appendix to my paper, and it is to the omission of the remarks contained therein that I wish to direct attention.

In the description of the male flower, Strasburger and I almost completely agree. It possesses two outer parts of the perianth, two inner parts, six stamens, which I believe to arise by branching from two primordial stamens, although Strasburger does not agree with me in this, and two carpels. The formula of the flower may be expressed thus:—



In the female flower I had very great difficulty in coming to a conclusion as to the value of the two outer parts, but the inner I concluded was a covering of the nucleus, an ovular integument, and not carpellary. There were only two ways of deciding what was the morphological significance of the two outer parts, either by comparison with the male flower, or by comparing them with the parts in the flowers of *Ephedra* and *Gnetum*. I applied to Dr. Hooker for specimens of these genera, and he has kindly promised to procure them for me. As Strasburger's material for the examination of *Ephedra* and *Gnetum* was imperfect, it is still of importance to examine both in detail. Being, therefore, obliged to fall back on comparison with the male flower (the study of the development alone not being sufficient for the purpose), I described the two outer parts as forming a perianth, although I could not feel certain that I was correct in so doing, and could not explain the occurrence of the short stalk under them, no such stalk existing in the male. On looking at Strasburger's figures of *Ephedra*, I at once saw that I had been in error in describing the outer parts as forming a perianth, and in the appendix stated that they were *carpellary*.

The formula would therefore be:—



The carpels, therefore, exist in both flowers; but whereas in the male they are anterior and posterior, in the female they are lateral. Kindly make this correction, because I do not think that after Strasburger's magnificent work, the Gymnospermous theory is for a moment tenable.

Should any correspondent be able to obtain specimens of *Ephedra* and *Gnetum* for me, I would be greatly obliged, as I am desirous of completing my paper on "Welwitschia" by a description of its embryogeny, as well as that of the other two genera. Specimens which have been put in absolute alcohol are by far the best for examination, but that, I fear, could only be obtained abroad with great difficulty.

Dublin, Dec. 27, 1872

W. R. MCNAB

Gauges for Ocean Rainfall

IN reply to Mr. Miller's letter on ocean rainfall, in *NATURE*, vol. vii. p. 123, I beg to acquaint your correspondent that I have endeavoured to meet the difficulties he mentions, by designing two forms of rain-gauge for use on board ship. One is of a cylindrical form, and composed of a collector and receiver, detachable from each other, and is suspended on gimbals in a frame or vena. The rainfall may be estimated either by a glass scale at the sides, or by emptying the contents into a graduated glass measure.

A description of this instrument as above designed appeared in the *Journal of the Scottish Meteorological Society* for January 1870, and was illustrated by diagrams.

The other form consists of the cylinder as above, divided into col-